

the planetary formation models, and astronomers had no reason to think the solar

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system unusual.

"But we now know that these other planetary systems don't look like the solar system at all," said <u>Frederic A. Rasio</u>, a theoretical astrophysicist and professor of physics and astronomy in Northwestern's <u>Weinberg College of Arts and Sciences</u>. He is senior author of the Science paper.

"The shapes of the exoplanets' orbits are elongated, not nice and circular. Planets are not where we expect them to be. Many giant planets similar to Jupiter, known as 'hot Jupiters,' are so close to the star they have orbits of mere days. Clearly we needed to start fresh in explaining planetary formation and this greater variety of planets we now see."

Using the wealth of exoplanet data collected during the last 15 years, Rasio and his colleagues have been working to understand planet formation in a much broader sense than was possible previously. Modeling an entire planetary system -- the varied physical phenomena associated with gas, gravity and grains of material, on such a variety of scales -- was a daunting challenge.

The work required very powerful computers. The researchers also had to judiciously decide what information was important and what was not, so as to speed up the calculations. They decided to follow the growth of planets, the gravitational interaction between planets, and the whole planetary system in its entire spatial extent. They chose not to follow the gas disk's fluid dynamics in fine detail, but rather more generally. As a result, they were able to run simulations spanning a planetary system's entire formation.

The simulations suggest that an average planetary system's origin is extremely dramatic. The gas disk that gives birth to the planets also pushes them mercilessly toward the central star, where they crowd together or are engulfed. Among the growing planets, there is cut-throat competition for gas, a chaotic process that produces a rich variety of planet masses.

Also, as the planets approach each other, they frequently lock into dynamical resonances that drive the orbits of all participants to be increasingly elongated. Such a gravitational embrace often results in a slingshot encounter that flings the planets elsewhere in the system; occasionally, one is ejected into deep space. Despite its best efforts to kill its offspring, the gas disk eventually is consumed and dissipates, and a young planetary system emerges.

"Such a turbulent history would seem to leave little room for the sedate solar system, and our simulations show exactly that," said Rasio. "Conditions must be just right for the solar system to emerge."

Too massive a gas disk, for example, and planet formation is an anarchic mess, producing "hot Jupiters" and noncircular orbits galore. Too low-mass a disk, and nothing bigger than Neptune -- an "ice giant" with only a small amount of gas -- will grow.

"We now better understand the process of planet formation and can explain the properties of the strange exoplanets we've observed," said Rasio. "We also know that the solar system is special and understand at some level what makes it special."

"The solar system had to be born under just the right conditions to become this quiet place we see. The vast majority of other planetary systems didn't have these special properties at birth and became something very different."

In addition to Rasio, other authors of the Science paper are Edward W. Thommes, an adjunct professor at the University of Guelph in Ontario, former postdoctoral fellow at Northwestern and lead author, and Soko Matsumura, a postdoctoral fellow at Northwestern.

The computer simulations were performed on a supercomputing cluster operated by Northwestern's Theoretical Astrophysics Group and partially funded by a Major Research Instrumentation grant from the <u>National Science Foundation</u> (NSF). Rasio's research group on exoplanets also is funded by a grant from the NSF Division of Astronomy.

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